

Atomic Fine Structure

Note, we will use the Gottfried normalization for \hat{L}

~~\hat{L}~~ where \hbar is factored out $\hat{L}_{\text{Gottfried}} = \frac{\hat{L}}{\hbar}$

The Hamiltonian for the Hydrogen atom is

$$\hat{H}_0 = \frac{\hat{p}^2}{2\mu} - \frac{e^2}{r}$$

$$\mu = \frac{m_e m_p}{m_e + m_p} = 0.9995 m_e = \text{reduced mass}$$

$$E_n^0 = -\frac{\alpha^2 \mu c^2}{2n^2} \quad n=1, 2, 3, \dots$$

$$\alpha = \frac{e^2}{\hbar c} = \frac{1}{137.04} = \text{fine structure constant}$$

also write

$$E_n^0 = -\frac{e^2}{2a_0 n^2}$$

$$a_0 = \frac{\hbar^2}{\mu e^2} = \text{Bohr radius} = 0.529 \text{ \AA}$$

$$E_1^0 = -13.6 \text{ eV}$$

Corrections to \hat{H}_0 .

1.) Relativistic effects

In reality the kinetic energy is $\sqrt{\mu^2 c^4 + p^2 c^2} - \mu c^2$

$$= \mu c^2 \left[1 + \frac{1}{2} \frac{p^2}{\mu^2 c^2} - \frac{1}{8} \frac{p^4}{\mu^4 c^4} + \dots \right] - \mu c^2$$

$$= \frac{p^2}{2\mu} - \frac{1}{8} \frac{p^4}{\mu^3 c^2}$$

$$\text{So } \boxed{V_{\text{rel}} \approx -\frac{1}{8} \frac{\mu^4}{p^4}}$$

2.) Spin-orbit coupling

Since the proton and electron both rotate about their center of mass, the electron sees a moving proton in its rest frame. This moving charge creates a magnetic field that the electron interacts with.

The electron magnetic moment gives an energy

$$-\vec{\mu} \cdot \vec{H} \quad \begin{array}{l} \vec{\mu} = \text{magnetic moment not reduced mass} \\ \vec{H} = \text{magnetic field not Hamiltonian} \end{array}$$

$$\vec{\mu} = \frac{e\hbar}{2mc} \vec{\sigma} \quad \vec{\sigma} = \text{Pauli spin matrix}$$

$$\vec{H} = -\frac{1}{c} \vec{v} \times \vec{E} = \frac{1}{ec} \vec{v} \times \hat{r} \frac{dV}{dr}$$

$$= -\frac{\hbar}{emc} \vec{L} \frac{1}{r} \frac{dV}{dr} \quad \text{with } V(r) = -\frac{e^2}{r}$$

from
Gottfried
normalization

We add the term

$$\frac{1}{4} \left(\frac{\hbar}{mc} \right)^2 \vec{L} \cdot \vec{\sigma} \frac{1}{r} \frac{dV}{dr}$$

which has an extra factor of $\frac{1}{2}$ multiplying it (called Thomas precession factor and very hard to derive)

This is called spin-orbit coupling

$$\text{so } \hat{H} = \hat{H}_0 + \hat{V}$$

$$\hat{H}_0 = \frac{\hat{p}^2}{2\mu} - \frac{e^2}{r} \quad \hat{V} = -\frac{1}{8} \frac{\hat{p}^4}{\mu^3 c^2} + \frac{1}{2} \left(\frac{\hbar}{mc} \right)^2 \hat{L} \cdot \hat{S} \frac{1}{r} \frac{dV}{dr}$$

$$\hat{S} = \vec{\sigma} / 2$$

The relativistic perturbation \hat{p}^4 is a scalar and commutes with all angular momenta and spin

Note $\hat{L} \cdot \hat{S} = \frac{\hat{L}_+ \hat{S}_- + \hat{L}_- \hat{S}_+}{2} + \hat{L}_z \hat{S}_z$

so $[\hat{L} \cdot \hat{S}, \hat{L}_z] = \frac{\hat{L}_+ \hat{S}_- + \hat{L}_- \hat{S}_+}{2} \neq 0$
 $[\hat{L} \cdot \hat{S}, \hat{S}_z] = \frac{\hat{L}_+ \hat{S}_- - \hat{L}_- \hat{S}_+}{2} \neq 0$

but $[\hat{L} \cdot \hat{S}, \hat{L}_z + \hat{S}_z] = 0$

so, a set of mutually commuting variables is

$\hat{J}^2, \hat{J}_z, \hat{L}^2, \hat{S}^2$

($\hat{J}^2 = \hat{L}^2 + 2\hat{L} \cdot \hat{S} + \hat{S}^2$ commutes with $\hat{L} \cdot \hat{S}$ too)

so label states by $|j, m; l, s = \frac{1}{2}\rangle$

We use spectroscopic notation to label the states

- $n l_j$ $n =$ principal quantum number
- $l =$ s, p, d, f
- $l=0 \quad l=1 \quad l=2 \quad l=3$
- $j =$ half odd integer

like $1S_{1/2} \quad 2S_{1/2} \quad 2P_{3/2} \quad 2P_{1/2}$ etc.

since $\hat{J}^2 = (\hat{L} + \hat{S})^2 = \hat{L}^2 + 2\hat{L} \cdot \hat{S} + \hat{S}^2$

so $\hat{L} \cdot \hat{S} = \frac{\hat{J}^2 - \hat{L}^2 - \hat{S}^2}{2}$ (a useful identity)

hence $\langle j, m; l, s = \frac{1}{2} | \hat{L} \cdot \hat{S} | j, m; l, s = \frac{1}{2} \rangle = \frac{j(j+1) - l(l+1) - \frac{3}{4}}{2}$

so $V_{so} = \frac{1}{4} \frac{\hbar^2}{m^2 c^2} [j(j+1) - l(l+1) - \frac{3}{4}] \frac{e^2}{r^3}$ when acting on these states

Now, since $j = l \pm \frac{1}{2}$ we have

$$\hat{V}_{so} = \frac{1}{4} \frac{\hbar^2 e^2}{m^2 c^2} \frac{1}{r^3} \left[(l \pm \frac{1}{2})(l + 1 \mp \frac{1}{2}) - l(l+1) - \frac{3}{4} \right]$$

$$= \frac{1}{4} \frac{\hbar^2 e^2}{m^2 c^2} \frac{1}{r^3} \begin{cases} +l & j=l+\frac{1}{2} \\ -l-1 & j=l-\frac{1}{2} \end{cases}$$

$$\langle n l j m | \hat{V}_{so} | n l j m \rangle = \frac{1}{4} \left(\frac{\hbar}{m c} \right)^2 e^2 \langle n l | \frac{1}{r^3} | n l \rangle * \begin{cases} l & j=l+\frac{1}{2} \\ -l-1 & j=l-\frac{1}{2} \end{cases}$$

The radial integral is known $\langle n l | \frac{1}{r^3} | n l \rangle = \frac{1}{a_0^3} \frac{1}{n^3 (l+1)(l+1/2)l}$ for $l \neq 0$

$$\hat{V}_{kin} = -\frac{1}{8} \frac{\hat{p}^4}{m^3 c^2} = - \left[\hat{H}_0 + \frac{e^2}{r} \right]^2 \cdot \frac{1}{2m c^2} \quad \text{another use full trick}$$

$$\langle n l j m | \hat{V}_{kin} | n l j m \rangle = - \langle n l j m | \left[\hat{H}_0 + \frac{e^2}{r} \right]^2 | n l j m \rangle \frac{1}{2m c^2}$$

$$= - \left[E_n^0{}^2 + 2E_n^0 e^2 \langle n l | \frac{1}{r} | n l \rangle + e^4 \langle n l | \frac{1}{r^2} | n l \rangle \right] \frac{1}{2m c^2}$$

These radial integrals are also known $\langle n l | \frac{1}{r} | n l \rangle = \frac{1}{a_0 n^2}$

$$\langle n l | \frac{1}{r^2} | n l \rangle = \frac{1}{a_0^2 n^3 (l+1)}$$

and $E_n^0 = -\frac{e^2}{2a_0 n^2}$ so we get

$$\langle n l j m | \hat{V}_{kin} | n l j m \rangle = -\frac{e^4}{4a_0^2} \left[\frac{1}{n^4} - \frac{4}{n^4} + \frac{4}{n^3(l+1)} \right] \frac{1}{2m c^2}$$

$$= E_n^0 \frac{e^2}{a_0} \frac{1}{n^2} \left[\frac{n}{2l+2} - \frac{3}{4} \right] \frac{1}{m c^2}$$

$$= E_n^0 \alpha^2 \frac{1}{n^2} \left[\frac{n}{2l+2} - \frac{3}{4} \right] \quad \alpha = \frac{e^2}{\hbar c}$$

similarly, we find

$$\begin{aligned} \langle n l j m | \hat{V}_{so} | n l j m \rangle &= \frac{1}{4} \frac{\hbar^2}{m^2 c^2} \frac{e^2}{a_0} \frac{\omega^2 e^4}{\hbar^4} \frac{1}{n^3} \frac{1}{(l+1)(l+\frac{1}{2})l} * \int_{-l-1}^{l+\frac{1}{2}} \int_{j=l+\frac{1}{2}}^{j=l+\frac{1}{2}} \\ &= -E_n^0 \alpha^2 \frac{1}{n} \frac{1}{2(l+1)(l+\frac{1}{2})l} \int_{-l-1}^{l+\frac{1}{2}} \int_{j=l+\frac{1}{2}}^{j=l+\frac{1}{2}} \end{aligned}$$

$j = l + \frac{1}{2}$ case

$$\begin{aligned} E_n &= E_n^0 \left[1 + \frac{\alpha^2}{n^2} \left[\frac{n}{j} - \frac{3}{4} - \frac{n}{(j+\frac{1}{2})2j} \right] \right] \\ &= E_n^0 \left[1 + \frac{\alpha^2}{n^2} \left[\frac{n}{j+\frac{1}{2}} - \frac{3}{4} \right] \right] \end{aligned}$$

$j = l - \frac{1}{2}$ case

$$\begin{aligned} E_n &= E_n^0 \left[1 + \frac{\alpha^2}{n^2} \left[\frac{n}{j+1} - \frac{3}{4} + \frac{n(j+\frac{3}{2})}{2(j+\frac{3}{2})(j+1)(j+\frac{1}{2})} \right] \right] \\ &= E_n^0 \left[1 + \frac{\alpha^2}{n^2} \left[\frac{n}{j+\frac{1}{2}} - \frac{3}{4} \right] \right] \end{aligned}$$

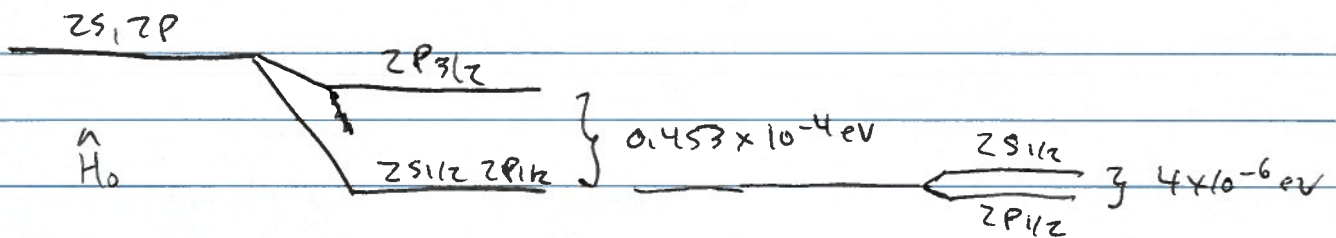
recall $E_n^0 < 0 \Rightarrow$ shift is negative (book has a typo)

Comments

1.) even though this shift wasn't calculated for $j = \frac{1}{2}, l = 0$, it holds for that case as well (need Dirac equation to prove it)

2.) The lowest excited states are $2P_{1/2}, 2S_{1/2}$

which are not split from each other since shift depends only on j
 experimentally they are split (called the Lamb shift)
 which can be understood only with quantum electrodynamics (field theory)



- 3.) By choosing the jm basis, we did not need to actually use degenerate perturbation theory because we found the Π subspace.

In next lecture, we will need the degenerate formalism.