Phase separation and the segregation principle in the infinite-$U$ spinless Falicov-Kimball model

J. K. Freericks

Department of Physics, Georgetown University, Washington, D.C. 20057

Ch. Gruber and N. Macris

Institut de Physique Théorique, Ecole Polytechnique Fédérale de Lausanne, PHB-Ecublens, CH-1015, Lausanne, Switzerland

(Received 26 January 1999)

The simplest statistical-mechanical model of crystalline formation (or alloy formation) that includes electronic degrees of freedom is solved exactly in the limit of large spatial dimensions and infinite interaction strength. The solutions contain both second-order phase transitions and first-order phase transitions (that involve phase separation or segregation) which are likely to illustrate the precursor physics behind the static charge-stripe ordering in cuprate systems. In addition, we find that the spinodal-decomposition temperature satisfies an approximate scaling law.

I. INTRODUCTION

The most fundamental problem in solid-state physics is to understand why elements (and most compounds) crystallize in ordered periodic structures, for this forms the basis of all of solid-state physics. While it is well known that the driving principle behind this ordering is a lowering of the ground-state energy of the material, and there has been significant progress with ab initio methods to predict the ground-state properties of these ordered phases in real materials, progress has been slow in illuminating the solution of the simplest model for crystal formation (that describes the statistical-mechanical mechanism behind the ordering of the electrons and ions on a periodic lattice). Furthermore, it is not understood what physical mechanisms are necessary for creating a crystallized state. This crystallization problem is ubiquitous; it also describes the statistical mechanics behind binary-alloy formation or phase separation since the two problems can be mapped onto each other (as described below), and it may also describe the physics behind charge-stripe formation in the cuprates.

The solution of a statistical-mechanical model for magnetic order has been known ever since Onsager solved the two-dimensional Ising model.1 Onsager’s solution produced a paradigm for understanding phase transitions in many different systems, and provided a textbook example of much of the theory behind modern critical phenomena. In fact, Lee and Yang2 modified the Ising model to consider the magnetic order in an external magnetic field, and mapped the problem onto a lattice gas, where the up-spins denoted sites occupied by ions, and the down-spins denoted empty sites. Onsager’s method of solution does not extend to the case of a finite magnetic field, so no exact results are known for the lattice gas, except in the case where the number of ions equals one-half the number of lattice sites, which corresponds to the zero-field case. These models of crystallization neglect the electronic degrees of freedom of the valence electrons, and hence are not directly applicable to real materials such as metals and alloys.

It turns out that the Ising model, and many other models for magnetism, simplify when they are examined in high dimensions. In fact, the critical behavior of the Ising model is determined by a static mean-field theory in four dimensions and higher. A similar situation is expected for electronic problems, except they remain nontrivial even in the infinite-dimensional limit.3,4 Metzner and Vollhardt showed that the electronic problem requires a dynamical mean-field theory for its solution in infinite dimensions. Furthermore, a wide range of evidence indicates that this dynamical mean-field theory provides a quantitative approximation to the solutions of correlated electron problems in three dimensions (at least if one is not too close to a critical point). In fact, it is precisely the nonuniversal properties (such as a transition temperature) that the dynamical mean-field theory determines accurately, and its solution provides a wealth of information on the qualitative behavior of the model studied.

Here we employ dynamical mean-field theory to produce an exact solution of the simplest crystallization problem which includes the electronic degrees of freedom.

This model is the spinless Falicov-Kimball model5 which consists of two kinds of particles: localized ions and itinerant (spinless) electrons. The localized ions ($w_i = 0$ or 1) occupy sites on a lattice in real space with an energy $E$, and the electrons can hop (with a hopping integral $-t^* [2 \sqrt{d}]$) between neighboring lattice sites. In addition, there is a screened Coulomb interaction $U$ between electrons and ions that occupy the same lattice site. Since the electrons do not interact with each other, the “spin” degree of freedom is unimportant, and is neglected. The Hamiltonian is

$$H = -\frac{t^*}{2\sqrt{d}} \sum_{\langle i,j \rangle} c_i^\dagger c_j + E \sum_i w_i + U \sum_i c_i^\dagger c_i w_i,$$

(1)

with $c_i^\dagger$ ($c_i$) the creation (annihilation) operator for electrons at site $i$, and $w_i$ denoting the ion occupancy at site $i$. We use $t^* = 1$ as the energy scale.

As it stands, the Falicov-Kimball model does not appear to be a many-body problem at all, since the ions are localized and do not move, which implies that the quantum-mechanical problem for the electrons can be solved by diagonalizing a single-particle problem of an electron moving in the potential determined by the given configuration of the
ions \( \{w_i\} \). The many-body problem aspects enter by taking an annealed average over all possible ion configurations with the chosen ion concentration. This produces long-range interactions between the ions, that can cause them to order or phase separate at low temperatures.

The Falicov-Kimball model can be viewed as a simplified approximation of a real material in a variety of ways. If the material has a single valence electron, and only one electronic band lies near the Fermi level, then the crystallization problem would correspond to the case where the electron and ion concentrations \( (\rho_e, \rho_i) \) are the same (which is called the neutral case), since one electron is donated by each ion. If, instead, there are many bands near the Fermi level, then one can map the combined bands into a single “effective” band which will have an electron filling determined by the average filling of the electrons in the most important band. In this case, each ion may donate only a fraction of an electron to the crystal, because the rest of the electron goes into other hybridized bands that lie close to the Fermi level. Hence one may find it useful also to consider non-neutral cases for the crystallization problem, where the electron and ion concentrations are not equal. This model can also be mapped onto the binary alloy problem, where a site occupied by an ion is screened Coulomb interaction is mapped to the difference in electronic degrees of freedom are introduced.

The Falicov-Kimball model is expected to be in the same universality class as the Ising model, but, because of the electronic degrees of freedom, one needs to solve the full statistical model to determine the “effective magnetic exchange parameters” between different lattice sites. The parameters can be extracted in a systematic expansion if the electronic kinetic energy (the hopping term) is taken as a perturbation, but such an analysis is only valid in the strong-coupling regime, and rapidly becomes problematic. It is precisely this complication that has frustrated attempts at finding an exact solution to the crystallization problem when electronic degrees of freedom are introduced.

The one-dimensional limit of the Falicov-Kimball model has also been extensively studied. Here there are no finite-temperature phase transitions, but the system can have phase transitions in the ground state. The first attempt at studying the one-dimensional Falicov-Kimball model proceeded along the lines of \textit{ab initio} band-structure calculations for real materials—a small number of candidate ion configurations were chosen for the ground state, and a restricted phase diagram was determined for all structures within the subset. The numerical solutions produced two conjectures: the first was a result for the case where \( \rho_e \neq 1 - \rho_i \), which stated that, if the screened Coulomb interaction \( U \) was large enough, then the system would segregate into an empty lattice (with no ions and all the electrons), and a full lattice (with all the ions and no electrons). The second was a generalization of the Peierls instability, which says that in the small-\( U \) limit the system will order in such a fashion that the ions produce a band structure that has a maximal gap at the Fermi level. This first conjecture (the segregation principle) was later proven to be true by Lemberger, while the second conjecture was shown to be false if the electron concentration was sufficiently far from half-filling. In that case, the system would phase separate between the empty lattice, and an optimally chosen ion structure that had the Fermi level lying in the gap.

The other limit that has been extensively studied is the large-dimensional limit, where Brandt and Mielsch provided the solution of the transition temperature as a function of \( U \) for the half-filled symmetric case. Their solution involves solving a coupled set of transcendental equations which display first- and second-order phase transitions. Freericks later showed that the model (on a hypercubic lattice) also displayed incommensurate order and segregation.

There are two kinds of lattices that are usually investigated in the large coordination-number limit: the hypercubic lattice, which is the generalization of the cubic lattice to large dimensions; and the Bethe lattice, which is a thermodynamic limit of the Cayley tree when the number of nearest neighbors becomes large. The noninteracting band structure for the hypercubic lattice produces a density of states that is a Gaussian \( \rho_\text{pert}(\varepsilon) = \exp(-\varepsilon^2/\sqrt{\pi}) \), while on the Bethe lattice the density of states is semicircular \( \rho_\text{Bethe}(\varepsilon) = \sqrt{4 - \varepsilon^2/(2 \pi)} \). The hypercubic density of states has an infinite bandwidth, but most of the weight lies within a range of \( \pm 2 \) about the origin. The Bethe lattice density of states has the same behavior as a three-dimensional system at the band edge (square-root behavior), but has no van Hove singularities in the interior of the band. Because both densities of states are nontrivial, the many-body problem maintains much of its rich behavior that arises from the competition between kinetic-energy effects and interaction-energy effects. In particular, the Falicov-Kimball model continues to have phase transitions in the large coordination number limit, but the transitions have mean-field theory exponents.

In this contribution, we examine what happens in the case when the Coulomb interaction becomes infinite \( U \to \infty \) (the attractive case is equivalent to this case through a particle-hole transformation of the electrons, which carries \( \rho_e \to 1 - \rho_i \)). In this case, the electrons avoid the sites of the lattice occupied by the ions, so the electron concentration varies from zero up to \( 1 - \rho_i \). We investigated the non-unit-density cases, where the electron concentration was restricted to \( 0 \leq \rho_e < 1 - \rho_i \). In Sec. II the formalism and results for calculations on the Bethe lattice are presented. In Sec. III, results for the hypercubic lattice are given and in Sec. IV we present our conclusions.
II. FORMALISM AND RESULTS
FOR THE BETHE LATTICE

In the thermodynamic limit, the local lattice Green’s function is defined to be
\[ G_n = G(i\omega_n) = \int_0^\beta d\tau e^{i\omega_n\tau} \frac{\text{Tr}(e^{-\beta(H - \mu N)T_x}c(\tau)c^\dagger(0))}{\text{Tr}(e^{-\beta(H - \mu N)})}, \]
where \( i\omega_n = i\pi T(2n + 1) \) is the fermionic Matsubara frequency, \( \beta = 1/T \) is the inverse temperature, \( \mu \) is the electron chemical potential, and \( T_x \) denotes \( \tau \) ordering. The angle brackets in Eq. (2) denote the sum over ionic configurations. The local Green’s function is determined by mapping onto an atomic problem in a time-dependent field, with the action
\[ S_{\text{atomic}} = \int_0^\beta d\tau d\tau' c^\dagger(\tau)G_n^{-1}(\tau - \tau')c(\tau') + U\int_0^\beta d\tau c^\dagger(\tau)c(\tau)w + Ew, \]
where \( w = 0 \) and 1 is the ion number for the atomic site, and \( G_n^{-1} \) is the mean-field or effective-medium Green’s function, which is determined self-consistently (as described below). The atomic Green’s function, with the action in Eq. (3), is computed to be
\[ G_n = \frac{1 - \rho_i}{G_0^{-1}(i\omega_n)} + \frac{\rho_i}{G_0^{-1}(i\omega_n) - U}, \]
with \( \rho_i \) the average ion density \( (w) \). On the other hand, the local lattice Green’s function satisfies
\[ G_n = \int_{-\infty}^{\infty} d\epsilon \frac{\rho(\epsilon)}{i\omega_n + \mu - \Sigma_n - \epsilon}, \]
where \( \rho(\epsilon) \) is the noninteracting density of states for the infinite lattice, and \( \Sigma_n \) is the self-energy. The self-consistency relation is that the self-energy \( \Sigma_n \) in Eq. (5) must coincide with the self-energy of the atomic problem, i.e.,
\[ \Sigma(i\omega_n) = G_0^{-1}(i\omega_n) - G_n^{-1}. \]

Equations (4), (5), and (6) constitute the mean-field theory for homogeneous phases. In the limit \( d \to \infty \), Eq. (6) is an exact equation for the lattice problem. We note that for periodic phases, if they exist, one needs to replace the atomic problem by a more complicated many-site problem.\(^{17}\)

These equations are complicated to solve analytically, but a simplification occurs for \( 0 \leq \rho_e \leq 1 - \rho_i \) in the limit \( U \to \infty \). Indeed, when \( U \) is large the spectrum of the Hamiltonian consists of two bands separated by a gap of order \( U \) for electron fillings that satisfy \( 0 \leq \rho_e \leq 1 - \rho_i \). In this case the chemical potential lies within the lower band, so that \( \mu \) is \( O(1) \). We note that \( G_0 \) is a function of \( \mu \), and therefore for any finite \( \mu \), \( 1/(G_0(i\omega_n)U) \to 0 \) as \( U \to \infty \). Then Eq. (4) becomes
\[ G_n = (1 - \rho_i)G_0(i\omega_n), \]
and substituting Eq. (7) into Eq. (6) and solving for the self-energy then yields
\[ \Sigma_n = -\frac{\rho_i}{G_n} \]
for the relation between the local self energy and the Green’s function. Hence, in the limits \( U \to \infty \) and \( d \to \infty \) the equations for the homogeneous phase reduce to Eqs. (5) and (8). In the case of the Bethe lattice, \( \rho_n(\epsilon) = \sqrt{4 - \epsilon^2/(2\pi)} \), for \(-2 \leq \epsilon \leq 2 \), so that the integral in Eq. (5) can be performed analytically:
\[ G_n = \frac{i\omega_n + \mu - \Sigma_n}{2} - \frac{1}{2}\sqrt{(i\omega_n + \mu - \Sigma_n)^2 - 4}. \]
Substituting the result from Eq. (8) into Eq. (9), and solving for \( G_n \), yields the exact result for the interacting Green’s function in the strongly correlated limit:
\[ G_n = \frac{i\omega_n + \mu}{2} - \frac{1}{2}\sqrt{(i\omega_n + \mu)^2 - 4(1 - \rho_i)}, \]
where the phase of the square root is chosen so that the Green’s function has the correct sign to its imaginary part. This form is identical to that of a noninteracting Green’s function [Eq. (9), with \( \Sigma_n = 0 \)], with a bandwidth narrowed from 4 to \( 4\sqrt{1 - \rho_i} \) and containing a spectral weight of \( 1 - \rho_i \) (since the remaining spectral weight is shifted to infinite energies). This is easiest seen from the interacting density of states, which satisfies\(^{18}\)
\[ \rho_n^{\text{int}}(\epsilon) = \frac{1}{2\pi}\sqrt{4(1 - \rho_i) - \epsilon^2}. \]
Note that in the infinite-interaction-strength limit, we have an analytic form for the Green’s functions, and do not need to solve transcendental equations iteratively as is normally done in the finite-U case.\(^{15}\) Furthermore, even though the Green’s function has the same form as a noninteracting Green’s function, the self-energy is nontrivial and does not correspond to a Fermi liquid.

This form for the Green’s function fits a rather simple physical picture. The electron avoids sites occupied by an ion when \( U \to \infty \), so the number of available sites is reduced by the fraction \( 1 - \rho_i \). This means, on average, that the number of nearest neighbors is reduced by the same factor, which reduces the bandwidth by \( \sqrt{1 - \rho_i} \). The total spectral weight is also reduced from 1 to \( 1 - \rho_i \), because the upper band (with \( \rho_i \) states) is located at infinite energy. What is surprising is that this “hand-waving” argument is exact for the Bethe lattice (we will see below it is a good approximation for the hypercubic lattice, but is not exact).

The interacting density of states is temperature independent\(^{19}\) in the local approximation, which means that we can examine the ground state at \( T = 0 \) to see if the system phase separates, or if the homogeneous phase is lowest in energy. The ground-state energy for an ion concentration \( \rho_i \) and an electron concentration \( \rho_e \) is
\[ E(\rho_e, \rho_i) = \int_{-\infty}^\mu d\epsilon \rho_n^{\text{int}}(\epsilon)\epsilon, \]
with \( \mu \) the chemical potential defined by
\[
\rho_e = \int_{-\infty}^{\mu} d\varepsilon \rho_B^{\text{int}}(\varepsilon),
\]
and \( \rho_B^{\text{int}} \) the interacting density of states. Substituting in the exact result from Eq. (11) yields
\[
E(\rho_+ \rho_0) = -\frac{4}{3\pi} (1 - \rho_0)^{3/2} \left[ 1 - \frac{\mu^2}{4(1 - \rho_0)} \right]^{3/2},
\]
and
\[
\rho_e = \frac{1 - \rho_0}{\pi} \cos^{-1} \left( \frac{-\mu}{2\sqrt{1 - \rho_0}} \right) + \frac{\mu}{2\sqrt{1 - \rho_0}} \sqrt{1 - \frac{\mu^2}{4(1 - \rho_0)}}.
\]
Using Eqs. (14) and (15), we will show that the mixture of the state with no ions and an electron filling \( \rho_e/(1 - \rho_e) \) with the state with all ions and no electrons has a lower energy than the homogeneous state, i.e.,
\[
E(\rho_e \rho_+)(1 - \rho_0)E\left( \frac{\rho_e}{1 - \rho_0}, 0 \right) + \rho_0 E(0,1).
\]
Moreover, from Eq. (16) we will deduce that the mixture corresponding to the right-hand side of Eq. (16) has a lower energy than any other mixture between homogeneous states. In other words,
\[
\alpha E(\rho_e', \rho_0') + (1 - \alpha) E(\rho_e'', \rho_0'')
> (1 - \rho_0) E\left( \frac{\rho_e}{1 - \rho_0}, 0 \right) + \rho_0 E(0,1),
\]
where \( 0 \leq \alpha \leq 1 \) and \( \rho_e = \alpha \rho_e' + (1 - \alpha) \rho_e'' \), \( \rho_0 = \alpha \rho_0' + (1 - \alpha) \rho_0'' \), \( 0 < \rho_e' < 1 - \rho_0' \), and \( 0 < \rho_e'' < 1 - \rho_0'' \). To obtain Eq. (16), we first note that \( E(0,1) = 0 \), and that the chemical potential \( \tilde{\mu} \) corresponding to an electron filling of \( \rho_e/(1 - \rho_e) \) and an ion filling of zero is \( \tilde{\mu} = \mu/(\sqrt{1 - \rho_e}) \), as can be seen from Eq. (15). Therefore, Eq. (14) yields
\[
(1 - \rho_0) E\left( \frac{\rho_e}{1 - \rho_0}, 0 \right) + \rho_0 E(0,1)
= -\frac{4}{3\pi} (1 - \rho_0)^{3/2} \left[ 1 - \frac{\mu^2}{4(1 - \rho_0)} \right]^{3/2}
= \frac{1}{\sqrt{1 - \rho_0}} E(\rho_e \rho_0) E(\rho_e \rho_0).
\]
which proves Eq. (16). The proof of Eq. (17) relies on an application of Eq. (16):
\[
\alpha E(\rho_e', \rho_0') + (1 - \alpha) E(\rho_e'', \rho_0'')
> \alpha \left[ (1 - \rho_0') E\left( \frac{\rho_e'}{1 - \rho_0'}, 0 \right) + \rho_0' E(0,1) \right]
+ (1 - \alpha) \left[ (1 - \rho_0'') E\left( \frac{\rho_e''}{1 - \rho_0''}, 0 \right) + \rho_0'' E(0,1) \right].
\]
The right-hand side of Eq. (19) is equal to
\[
(1 - \rho_0) \left[ \frac{\alpha \rho_e' + (1 - \alpha) \rho_e''}{1 - \rho_0} \right] E\left( \frac{\rho_e'}{1 - \rho_0'}, 0 \right)
+ \frac{(1 - \alpha) \rho_e''}{1 - \rho_0''} E\left( \frac{\rho_e''}{1 - \rho_0''}, 0 \right) + \rho_0 E(0,1).
\]
On the other hand, \( E(\rho_0, 0) \) is a convex function of \( \rho_0 \), so the term inside the brackets in Eq. (20) is greater than \( E(\rho_0, 0) \), which yields Eq. (17). We remark that the convexity of \( E(\rho_0, 0) \) is obvious from the fact that the free-electron system cannot phase separate. Formally, it can be seen as follows: differentiating Eqs. (14) and (15) with respect to \( \rho_e \) gives \( E'(\rho_e, 0) = \mu \rho_e \rho_0 / \rho_e \) and \( 1 = p_0(\mu) \delta \mu / \delta \rho_e \). Thus \( E'(\rho_e, 0) \) is \( \mu \) and \( E''(\rho_e, 0) = \delta \mu / \delta \epsilon = 1 / \rho_0(p_0(\mu)) > 0 \).
Our interest now is to determine the finite-temperature phase diagram of the infinite-\( U \) Falicov-Kimball model, since we know the system always phase separates at low temperature (although we have not yet ruled out the possibility of charge-density-wave phases being lower in energy than the phase-separated ground state). The first step is to evaluate the conduction-electron charge-density-wave susceptibility. Most work on the Bethe lattice has examined only antiferromagnetic or uniform order. Recent work has shown the existence of a period-three phase stabilized on the infinite-dimensional Bethe lattice at zero temperature. This ordering (and more complicated ones) can be associated with a one-dimensional momentum vector \( 0 \leq k \leq \pi \) that represents the modulation of the ordered phase along the levels of the Bethe lattice. The momentum dependence of the dressed susceptibility enters only through the momentum dependence of the bare susceptibility, because the vertex function is local in the infinite-dimensional limit. This allows us to simply take the \( U \to \infty \) of the Brandt-Mielsch result, which gives
\[
1 = \rho_0 \sum_{n=0}^{\infty} \frac{G_n^2 + \chi_n^0(X)}{G_n^2 + p_0\chi_n^0(X)},
\]
with \( X = \cos k \) being the parameter that determines the modulation of the charge-density wave over the Bethe lattice, and \( \chi_n^0(X) \) the corresponding bare susceptibility. We do not provide the general formula for all possible charge-density waves here. Rather, we present the three simplifying cases for the susceptibility on the Bethe lattice: (i) the local susceptibility, \( \chi_n^0(\text{local}) = -G_n^2 \); (ii) the \( X = -1 \) “antiferromagnetic” susceptibility, where
\[
\chi_n^0(-1) = -\frac{G_n}{i \omega_n + \mu - \Sigma_n};
\]
and (iii) the \( X = 1 \) uniform susceptibility, where
\[
\chi_n^0(1) = -\frac{\partial G_n}{\partial \mu} = -\frac{G_n}{\sqrt{(i \omega_n + \mu - \Sigma_n)^2 - 4}}.
\]
The local susceptibility never has a transition, because the numerator of Eq. (21) vanishes. The condition for an “anti-ferromagnetic” charge-density wave becomes

$$1 = \rho_i \sum_{n=-\infty}^{\infty} \frac{G_n}{i \omega_n + \mu},$$  \hspace{1cm} (24)

after substituting in the infinite-$U$ form for the self-energy, and using the quadratic equation $G_n^2 - (i \omega_n + \mu)G_n + 1 - \rho_i = 0$ that the interacting Green’s function satisfies. Now, substituting the integral form for $G_n$,

$$G_n = (1 - \rho_i) \int_{-\infty}^{\infty} \frac{\rho_B(e)}{i \omega_n + \mu - \sqrt{1 - \rho_i} e} \, de,$$  \hspace{1cm} (25)

into Eq. (24), and performing the sum over Matsubara frequencies yields the final integral form for $T_c$

$$1 = \rho_i \frac{\sqrt{1 - \rho_i}}{2T} \pi \int_{-\sqrt{1 - \rho_i}}^{\sqrt{1 - \rho_i}} \frac{z}{\sqrt{1 - \rho_i}} \tan \beta z \frac{\beta z}{\cosh^2 \frac{\beta \mu}{2} \left( 1 - \tanh \frac{\beta \mu}{2} \tan \frac{\beta z}{2} \right)} \, dz,$$  \hspace{1cm} (26)

(see the Appendix). But this integrand is positive for all $z$, so the right-hand side is always less than zero, and there is no “anti-ferromagnetic” $T_c$. The staggered charge-density-wave order has been found near half-filling $\rho_i = \rho_s = \frac{1}{2}$, when the lowest-order exchange for finite $U$ is included, but can only occur at $T=0$ and $\rho_i = \rho_s = \frac{1}{2}$ when $U=\infty$.

The uniform susceptibility case is analyzed as follows: First the uniform susceptibility from Eq. (23) is substituted into Eq. (21), and the square root is eliminated by using the exact form for the interacting Green’s function in Eq. (9). Next the self-energy is replaced by its exact form from Eq. (8), and the quadratic equation for $G_n$ is used to simplify the $T_c$ equation to

$$1 = \rho_i \sum_{n=-\infty}^{\infty} \left[ 1 + \frac{1 - \rho_i}{(i \omega_n + \mu)G_n - 2(1 - \rho_i)} \right].$$  \hspace{1cm} (27)

Now the interacting form for $G_n$ from Eq. (10), is substituted into Eq. (27), and the results simplified to yield

$$1 = \rho_i \sum_{n=-\infty}^{\infty} \frac{(i \omega_n + \mu)G_n - 2(1 - \rho_i)}{(i \omega_n + \mu)^2 - 4(1 - \rho_i)}.$$

The final step is to substitute in the integral form for $G_n$ from Eq. (25) and perform the summation over Matsubara frequencies (see the Appendix). After making a trigonometric substitution, the transcendental equation for $T_c$ becomes

$$1 = \rho_i \frac{\sqrt{1 - \rho_i}}{2 \pi T} \int_0^\pi d \theta \cos \theta \tanh \frac{\beta}{2} (2 \sqrt{1 - \rho_i} \cos \theta - \mu).$$  \hspace{1cm} (29)

We do not discuss any of the other periodic cases here, because the numerics involved is cumbersome. But we expect the Bethe lattice to have similar behavior as the hypercubic lattice, where the transition always went into the uniform charge-density wave, signifying a phase-separation transition. Details of the other periodic phases will be reported in a future publication.

The results for the transition temperature for the uniform charge-density wave are presented in Fig. 1(a). We choose nine different ion concentrations ranging from 0.1 to 0.9 in steps of 0.1. The electron density then varies from 0 to $1 - \rho_i$ for each case. As can be seen in the figure, the maximal transition temperature is about $0.12 \rho_s^* \rho_e$, and it occurs at half-filling of the lower band $\rho_e^* = (1 - \rho_i)/2$ with $\rho_s^* \approx 0.65 \rho_e^*$. Coincidentally, this maximal transition temperature is nearly identical to the maximal $T_c$ to charge-density-wave order at $\rho_e = \rho_s = \frac{1}{2}$ when evaluated as a function of the interaction strength $U$). Since $T_c \approx 1$, we expand Eq. (29) for small $T$ by replacing the $\tanh x$ by $\frac{x}{\sinh x}$ to find

$$T_c \approx \frac{\rho_i}{\sqrt{1 - \rho_i}} \frac{\frac{\beta^2 \mu}{4} (1 - \rho_i)}{1 - \rho_i}.$$  \hspace{1cm} (30)

Since the chemical potential will scale with $\sqrt{1 - \rho_i}$ for the same relative electron filling in the lower band $[\rho_e^*/(1 - \rho_i)]$, as shown in Eq. (15), this form motivates a scaling plot of $T_c \rho_e^*/(1 - \rho_i)$ versus $\rho_e^*/(1 - \rho_i)$, which appears in Fig. 1(b). As can be seen there, the data nearly collapse on top of each other for $T_c$ which is usually a nonuniversal quantity. In fact, the variation in $T_c$ is less than 10% for all different cases.

The susceptibility analysis shows that the system orders in a uniform charge-density wave, which indicates that the system will phase separate (or segregate) into two regions: one with a higher concentration of electrons and one with a lower concentration (as we already showed at $T=0$). Such a phase separation is usually associated with a first-order phase transition, rather than a second-order transition. Hence it is important to perform a Maxwell construction of the free energy that includes mixtures of two states with different electron and ion concentrations such that $\rho_e^* = \alpha \rho_e' + (1 - \alpha) \rho_e''$ and $\rho_i^* = \alpha \rho_i' + (1 - \alpha) \rho_i''$, and that the free energy of the mixture $F(\rho_e', \rho_i'; \rho_e'', \rho_i'') = \alpha F(\rho_e', \rho_i') + (1 - \alpha) F(\rho_e'', \rho_i'')$ is lower in energy than the pure-phase free energy $F(\rho_e, \rho_i)$. The second-order phase transition is the spinodal-decomposition temperature, below which the free energy becomes locally unstable in the region of $(\rho_e, \rho_i)$; in most cases the global free energy is minimized by the Maxwell construction at a temperature above this spinodal-decomposition temperature. The spinodal-decomposition temperature marks the lowest temperature that the system can be supercooled before it must undergo a phase transition.

We can calculate the free energy $F(\rho_e, \rho_i)$ for a homogeneous phase with electron filling $\rho_e$ and ion concentration $\rho_i$ in two equivalent ways. The first method is from Brandt and Mielsch, and expresses the free energy in terms of a summation over Matsubara frequencies as follows:
For the hypercubic lattice evaluated in Sec. III, we employ
First a coarse grid is established for
The numerical minimization proceeds in four phases: (i) First a coarse grid is established for \( \rho'_i \) and \( \rho''_i \), and the free energy is minimized over this grid [the electron fillings are determined by the constraints that the chemical potential is the same in region 1 and region 2 and that \( \rho_e = a \rho'_i + (1 - a) \rho''_i \), with \( a \) already determined from \( \rho_i = a \rho'_i + (1 - a) \rho''_i \)]. (ii) The filling \( \rho''_i \) is fixed at its coarse-grid minimal value, and \( \rho'_i \) is varied on a finer grid to determine the new minimum. (iii) \( \rho'_i \) is fixed at the new minimum, and \( \rho''_i \) is now varied on a fine grid to yield a new minimal \( \rho''_i \). (iv) \( \rho'_i \) and \( \rho''_i \) are varied together on the same fine grid to determine the final minimization of the Maxwell construction. We found that the minimal values of \( \rho'_i \) and \( \rho''_i \) rarely changed in step (iv), confirming the convergence of this method.

We plot our results in Fig. 2. The first case considered in Fig. 2(a) is the case of relative half-filling \( \rho_e = (1 - \rho_i)/2 \). In this case the chemical potential is always at zero, and the relative electron filling remains unchanged for all \( \rho_i \). The solid line is the first-order transition line, and the dotted line is the spinodal-decomposition temperature. The horizontal distance between the solid lines at a fixed temperature is a measure of the order parameter \( \rho'_i - \rho''_i \). Notice how the

where \( \rho^{int}(\epsilon) = \sqrt{4(1 - \rho_i) - \epsilon^2/2\pi} \) is the interacting density of states for the Bethe lattice. We find that both forms (31) and (32) are numerically equal to each other, but are unable to show this result analytically. Since the interacting density of states is known for the Bethe lattice, we use the

\[
F(\rho_e, \rho_i) = -T \ln \frac{1 + e^\beta \mu}{1 - \rho_i} + \int_{-\infty}^{\infty} d\epsilon \rho(\epsilon) T \sum_{n=-\infty}^{\infty} \ln \left[ \frac{i \omega_n + \mu}{(1 - \rho_i)(i \omega_n + \mu - \sum_{n'} - \epsilon)} \right] \\
+ \mu \rho_e - \left( T \ln \frac{\rho_i}{1 + \rho_i} + T \ln (1 + e^\beta \mu) + T \sum_{n=-\infty}^{\infty} \ln \left[ \frac{1 - \rho_i}{(i \omega_n + \mu - \sum_{n'}'} \right] \right) \rho_i. 
\]

Similarly, we can evaluate the free energy in the same fashion as Falicov and Kimball\(^5\) did:

\[
F(\rho_e, \rho_i) = T \int_{-\infty}^{\infty} d\epsilon \rho^{int}(\epsilon) \ln \left[ \frac{1}{1 + e^{-\beta (\epsilon - \mu)}} \right] + T[\rho_i \ln \rho_i + (1 - \rho_i) \ln (1 - \rho_i)] + \mu \rho_e, 
\]

FIG. 1. Second-order transition temperature on the Bethe lattice (corresponding to spinodal decomposition). (a) Transition temperature plotted as a function of electron filling. (b) Transition temperature plotted on a scaling curve as a function of relative electron filling.

FIG. 2. Transition temperature to phase separation on the Bethe lattice. (a) The case of relative half-filling \( \rho_e = (1 - \rho_i)/2 \). The solid line is the first-order transition temperature, and the dotted line is the spinodal decomposition temperature. Notice how these two curves meet at the maximum where the first-order transition becomes second order. (b) The case near relative quarter-filling (as described in the text). Notice how the shape of the curve differs from (a) near \( \rho_i = 1 \). This is because the electron filling becomes exponentially small once the chemical potential lies outside of the interacting bandwidth.
first-order transition temperature is always close to the spinodal-decomposition temperature, but that the difference becomes largest at concentrations close to zero and 1. Note further how the two curves meet at the maximum (as they must) where the first-order transition disappears and becomes a second-order transition at a classical critical point. In Fig. 2(b) we plot the phase diagram for the case with \( \rho_i = 0.65 \) and \( \rho_e = (1 - \rho_i)/4 \). In this case the chemical potential changes as a function of temperature, and as \( T \to 0 \) the system is phase separating into regions with ion densities close to zero and 1, we find that the chemical potential will lie outside of the bandwidth of the interacting density of states as \( \rho_i \to 1 \) because the bandwidth \( (4\sqrt{1 - \rho_i}) \) becomes narrowed to zero. In that case, the electron density approaches zero exponentially fast, which is why the spinodal-decomposition temperature approaches zero so rapidly in that regime. For this reason, we find that the relative electron filling is not a constant in this phase diagram, as it approaches zero exponentially fast near \( \rho_i = 1 \) and it is somewhat larger than quarter filled near \( \rho_i = 0 \). The two phase diagrams in Figs. 2(a) and 2(b) look similar in the low-density regime, however. This may imply that there is an analogous scaling regime for the first-order \( T_c \), but it does not look like there would be a universal curve for the region close to \( \rho_i = 1 \). The numerical effort required to perform the free-energy analysis is significant, so a thorough analysis of any possible scaling forms for \( T_c \) was not performed.

### III. RESULTS FOR THE HYPERCUBIC LATTICE

The formalism for the hypercubic lattice is essentially unchanged from the Bethe lattice. The main differences are that the integrals can no longer be performed analytically, which requires that results be worked out numerically, and requires more computational effort. The basic framework in Eqs. (2)–(8) is identical as above, except now the noninteracting density of states is a Gaussian for the hypercubic lattice. The integral for the local Green’s function is no longer elementary, and so one needs to solve the problem iteratively as was done previously for the Falicov-Kimball model: (i) first the self-energy is set equal to zero; (ii) next the local Green’s function is determined from Eq. (5); (iii) then the self-energy is determined from Eq. (8); (iv) then steps (ii) and (iii) are repeated until the equations converge. We can compare the results of the interacting Green’s function to the form found before for the Bethe lattice, by approximating the interacting density of states in the same fashion as before: we narrow the Gaussian by the factor \( \sqrt{1 - \rho_i} \), and have a total weight of \( 1 - \rho_i \) in the density of states. When we compare the Green’s function along the imaginary axis at half-filling in Fig. 3(a), we find that that approximation works well at high energies, but begins to fail near zero frequency (we chose \( \rho_e = \frac{1}{6}, \rho_i = \frac{2}{3}, \) and \( T = 0.1 \)). The solid line is the exact result, and the dotted line is the approximate (band-narrowed) result. The infinite-\( U \) Green’s function on the hypercubic lattice is more complicated than on the Bethe lattice, and the simple form that describes it for the Bethe lattice no longer holds. This is the main reason why the hypercubic lattice is more complicated to deal with than the Bethe lattice. To see this more fully, we examine the interacting density of states in Fig. 3(b). The interacting density of states is determined by solving the same equations for the Green’s function, but this time on the real axis, rather than the imaginary axis. Notice how the band-narrowed form \( \sqrt{1 - \rho_i} \exp(-\mu \vert 1 - \rho_i \vert) / \sqrt{2} \) (dotted line) is a reasonable approximation to the interacting density of states (solid line) but that it is too narrow, and it overestimates the peak height.

Since we do not know an analytic form for the interacting density of states, we cannot perform the same kind of analysis that we did before at zero temperature to see if the system is phase separated. But we can examine the finite-temperature phase diagrams in the same manner. The susceptibility diverges whenever Eq. (21) is satisfied. The bare susceptibility now takes the form

\[
\chi_n(X) = -\frac{1}{2\pi} \int_{-\infty}^{\infty} \rho(y) \frac{1}{i\omega_n + \mu - \Sigma_n - y} \int_{-\infty}^{\infty} dz \rho(z) \frac{1}{i\omega_n + \mu - \Sigma_n - yX - z\sqrt{1 - X^2}},
\]  

(33)
where $X = \lim_{d \to \infty} \frac{\sum_{j=1}^{d} \cos(k_j)}{d}$ for the ordering wave vector $\mathbf{k}$. The bare susceptibility continues to assume a simple form for the same three cases: (i) $X = 0$, the local susceptibility where $\chi_0^{(0)}(0) = -G_n^2$; (ii) $X = -1$, the ‘antiferromagnetic’ susceptibility, where $\chi_0^{(0)}(-1) = -G_n/(i\omega_n + \mu - \Sigma_n)$; and (iii) $X = 1$, the uniform susceptibility, where

$$
\chi_0^{(1)}(1) = \frac{\partial G_n}{\partial \mu} = 2[1 - (i\omega_n + \mu - \Sigma_n)G_n].
$$

If we try to approximate the transition temperature by substituting in the approximate form we derived for $G_n$ by assuming the interacting density of states has the same shape, but is band narrowed, we find that the $T_c$’s generated are not accurate at all. Hence the simple band-narrowing approximation works reasonably well for the Green’s function, but is poor for the susceptibility.

Instead, we simply solve for the transition temperatures numerically. We find in every case that we examined that the transition temperature is always highest for $X = 1$, and vanishes for all $X \leq 0$. This is shown in Fig. 4 for the cases of $\rho_e = \frac{1}{5}, \frac{1}{7}, \frac{1}{9}, \frac{1}{11}, \frac{1}{13}$, and $\frac{1}{15}$, and $\rho_i = \frac{1}{7}$, which ranges from relative half-filling to the low-density regime. We plot $T_c(X)$, and see that the system always favors the uniform charge-density wave, signifying that the system wants to phase separate. We calculate the spinodal-decomposition temperature for phase separation by finding the temperature at which the uniform susceptibility diverges as a function of $\rho_e$ and $\rho_i$. These temperatures are plotted in Fig. 5. This plot looks similar to what we found above for the Bethe lattice, so we try the same scaling form in Fig. 5(b), plotting $T_c(1 - \rho_e, \rho_i)$ versus $\rho_e/(1 - \rho_i)$. Once again we see a data collapse, but the spread in the $T_c$ ’s is somewhat larger than that seen in the Bethe lattice.

Finally, we calculate the full phase diagram for the case of relative half-filling $\rho_e/(1 - \rho_i)$, because they all satisfy a similar functional form. On the Bethe lattice we found that the system always phase separated at $T = 0$ to states where the electrons all moved to one part of the lattice, and the ions moved to the other part. The spinodal-decomposition temperature for segregation solved a simple transcendental equation, which we showed collapsed onto a scaling curve. In addition, we solved for the first-order transition temperature for a select number of cases, and discovered that the first-order transition usually occurred quite close to the spinodal-decomposition temperature. On the hypercubic lattice, we found similar results, but had to carry the analysis out numerically for all cases considered. We were able to show explicitly that phase separation precluded incommensurate (or commensurate) charge-density-wave order for the hypercubic lattice.

These results show that when the screened Coulomb interaction is large, or in the alloy picture, when the $A$ ions are extremely different from the $B$ ions, then the system will segregate at low temperatures. This proves the segregation principle for the infinite-dimensional limit, and leads us to believe that it holds for all dimensions (since it has also been demonstrated in one dimension). Future problems to be investigated include a study of the case $\rho_e + \rho_i = 1$, as well as finite Coulomb interaction $U$. In the unit-density case, we expect charge-density-wave order to be more prevalent, per-
haps precluding the segregated phase for all $U$. When the strength of the Coulomb interaction is reduced, we expect the segregated phase to disappear gradually, and be taken over by other phase-separated or charge-density-wave ordered phases.

It is possible that the phenomena described here incorporate the driving principle behind the stripe formation. There is in many strongly correlated models, and we believe that it is the desire for the system to phase separate, and the long-range Coulomb interaction, which would prevent the electrons from completely separating from the ions. We want to emphasize that we do not see striped phases in the solutions presented here, but rather we see phase separation, in the sense of the segregation principle, for the "simplified" $t$-$J$ model at $J=0$. This phase separation is ubiquitous in many strongly correlated models, and we believe that it is the driving principle behind the stripe formation. There is evidence for alternative points of view, however. White and Scalapino showed that the Hubbard model on a ladder displays charge-stripe order even without long-range Coulomb interaction. This order arises from the correlation of the spins and holes, and a desire to reduce the frustration induced by the hole motion. In their picture, the stripe ordering arises completely from a model that includes no long-range forces. Nevertheless, it is our belief that the phase separation exhibited here will be an important element of a complete description of the charge-stripe order in the cuprates and nickelates, because it must occur if $U$ becomes large enough.

ACKNOWLEDGMENTS

J.K.F. acknowledges support of this work from the Office of Naval Research Young Investigator Program.

\[ A_1 \]

Employing Eq. (29) involves summations over Matsubara frequencies, which are performed with the help of the identity

\[ \frac{1}{e^{\beta x} - 1} = \sum_{n=-\infty}^{\infty} \frac{1}{i\omega_n + x}, \]  

(A1)

for any real number $x$. Using Eq. (25) and the change of variables $\epsilon \rightarrow \epsilon / \sqrt{1 - \rho}$, yields

\[ \frac{G_n}{i\omega_n + \mu} = \sqrt{1 - \rho_i} \int_{-\infty}^{\infty} \frac{d\epsilon}{\rho_B} \left( \frac{\epsilon}{\sqrt{1 - \rho_i}} \right) \left[ \frac{1}{i\omega_n + \mu - \epsilon} - \frac{1}{i\omega_n + \mu + \epsilon} \right]. \]  

(A2)

Employing Eq. (A1) then shows that

\[ \sum_{n=-\infty}^{\infty} \frac{G_n}{i\omega_n + \mu} = \frac{2}{\beta} \sqrt{1 - \rho_i} \int_{-\infty}^{\infty} \frac{d\epsilon}{\rho_B} \left( \frac{\epsilon}{\sqrt{1 - \rho_i}} \right) \left[ \tanh \frac{\beta(\mu - \epsilon)}{2} - \tanh \frac{\beta \mu}{2} \right]. \]  

(A3)

Equation (26) then follows from the trigonometric identity

\[ \tanh \frac{\beta \mu}{2} - \tanh \frac{\beta \mu}{2} = \frac{\tan \mu}{\cosh^2 \frac{\beta \mu}{2} \left( 1 - \tanh \frac{\beta \mu}{2} \right) \tanh \frac{\beta \mu}{2}}. \]  

(A4)
The derivation of Eq. (29) is more involved, but proceeds along the same lines. Using the integral representation in Eq. (25) for \( G_n \), performing a decomposition into simple fractions, and then using the identity in Eq. (A1) produces both

\[
\sum_{n=-\infty}^{\infty} \frac{(i\omega_n + \mu)G_n}{(i\omega_n + \mu)^2 - 4(1 - \rho_i)} = \frac{\beta}{4\sqrt{1 - \rho_i}} \tanh \frac{\beta(\mu - 2\sqrt{1 - \rho_i})}{2} \int_{-\infty}^{\infty} d\epsilon \frac{\rho_B(\epsilon)}{2 - \epsilon} - \frac{\beta}{4\sqrt{1 - \rho_i}} \tanh \frac{\beta(\mu + 2\sqrt{1 - \rho_i})}{2} \int_{-\infty}^{\infty} d\epsilon \frac{\rho_B(\epsilon)}{2 + \epsilon}
\]

and

\[
\sum_{n=-\infty}^{\infty} \frac{2(1 - \rho_i)}{(i\omega_n + \mu)^2 - 4(1 - \rho_i)} = \frac{\beta}{4\sqrt{1 - \rho_i}} \left[ \tanh \frac{\beta(\mu - 2\sqrt{1 - \rho_i})}{2} - \tanh \frac{\beta(\mu + 2\sqrt{1 - \rho_i})}{2} \right].
\]

Now we use the fact that the integrals for the noninteracting Green’s function are trivial,

\[
\int_{-\infty}^{\infty} d\epsilon \frac{\rho_B(\epsilon)}{2 - \epsilon} = \int_{-\infty}^{\infty} d\epsilon \frac{\rho_B(\epsilon)}{2 + \epsilon} = 1.
\]

and subtract Eq. (A5) from Eq. (A6) to obtain

\[
\sum_{n=-\infty}^{\infty} \frac{(i\omega_n + \mu)G_n - 2(1 - \rho_i)}{(i\omega_n + \mu)^2 - 4(1 - \rho_i)} = \frac{\beta}{4\sqrt{1 - \rho_i}} \int_{-2}^{2} d\epsilon \frac{\epsilon}{\pi\sqrt{4 - \epsilon^2}} \tanh \frac{\beta(\epsilon \sqrt{1 - \rho_i} - \mu)}{2}.
\]

Equation (29) then follows from the change of variables \( \epsilon = 2\cos \theta \).